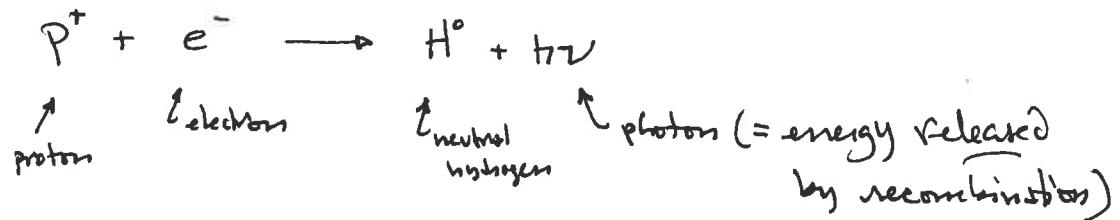


# The early universe

recall the process of recombination that occurred about 400,000 yrs after the big bang: led to the Universe becoming transparent to CMB photons

can write recombination as:



\* this reaction is exothermic

but it can run in reverse (endothermically)  
as long as the typical photon has enough  
energy (13.6 eV) to break apart  
the  $\text{H}^+$  atom

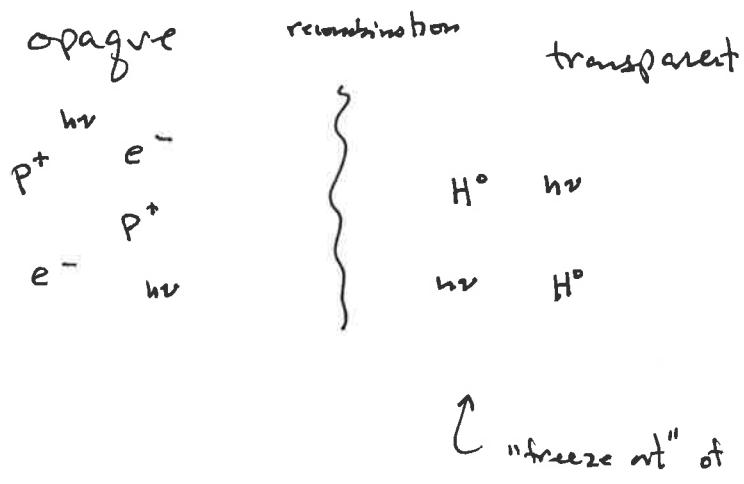


This reaction runs both ways BVT as the universe cools & expands it reaches a temperature where  $\Delta H^\circ$  isn't any longer high enough to drive the endothermic reaction, it then only runs in one direction & lots of  $H^\circ$  is made while  $P^\circ$  &  $e^\circ$  disappears. The  $H^\circ$  "freezes out".

Note:  $T_{\text{MB}} = T_{\text{MB},0} \alpha^{-1}$  tells us how big  $\alpha$  was when this freeze out occurred; if we have a model of  $\alpha(t)$  then we can know the time too

Note also that the redshift,  $z$ , corresponding to any scale factor is easily computed from  $1+z = 1/a$

$\leftarrow$  redshift, temperature  
time, scale factor  $\rightarrow$



Analogous situation w/ nuclear reactions (note:  $p^+ + e^- \rightarrow H^0 + h\nu$  is governed by electromagnetism)

e.g.



↑ deuterium: "fate of hydrogen"

\* reaction goes both ways early in the Universe, when the temperature is high

but when it "falls" to  $T \sim 10^9 \text{ K}$  it only goes to the right (photons no longer have the required 2.2 MeV to knock apart a deuterium nucleus)

At that point protons & neutrons get "used up" to make deuterium

but, the amount of deuterium that's made depends on:

1. how fast the universe is expanding & cooling (reducing the  $p + n \rightarrow D + \gamma$  rate as time goes on) And
2. how much of the D gets used up in other reactions

Fig 24.9 shows the time evolution of the various particles & simple nuclei that are created &/or destroyed in the various nuclear reactions described on p. 566-567

\* The fact that the actual, observed abundances of  $^4\text{He}$ ,  $^3\text{He}$ , D, &  $^7\text{Li}$  (etc.) agree with the predictions of this model is another very strong, independent piece of evidence for the hot big bang model of cosmology (along w/ the darkness of the night sky, the Hubble law, & the CMB).

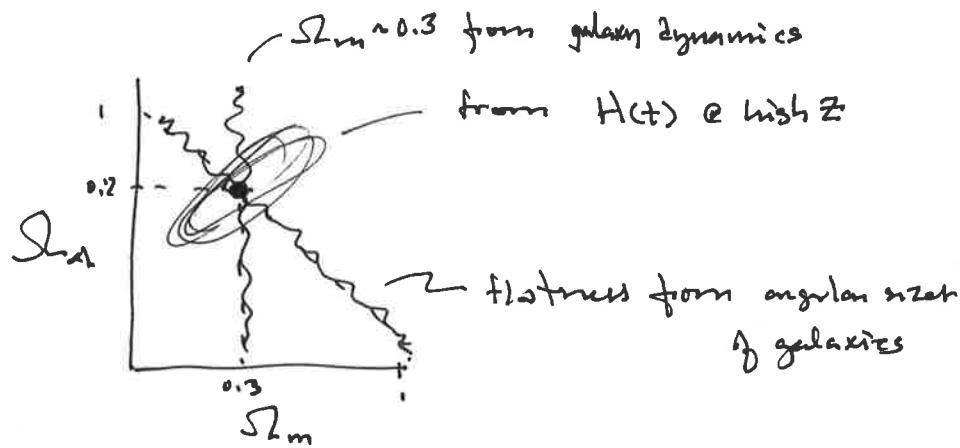
Now, the book doesn't state this explicitly but the abundance of these nuclei depends on the baryon density of the universe (as a function of time)  $\rightarrow$  higher density of baryons ( $p + n$ ) the more D is produced (& eventually He)

\* see ~~other~~ figure posted on website  $\rightarrow$  observed He, D, Li abundances constrain  $S_{L,b} \approx 0.04$  (where  $b = H_0/100 \text{ km/s/Mpc} \approx 0.7$ ) which is consistent w/ census of starlight + hot gas in galaxy clusters...

... and much less than dynamical determinations of  $\Omega_m \approx 0.3$   
 (confirming that  $\Omega_m \approx 0.26$  is <sup>non-</sup>"baryonic")

OK ... can thus summarize the consensus

model :



\* see photo of black board for more

note: regions of collapse  
 vs. expansion forever

note also — consensus model had 3 independent pieces of evidence  $\Rightarrow$  even if you take one away you still have  $\Omega_1 = 0.7$ ,  $\Omega_m = 0.3$

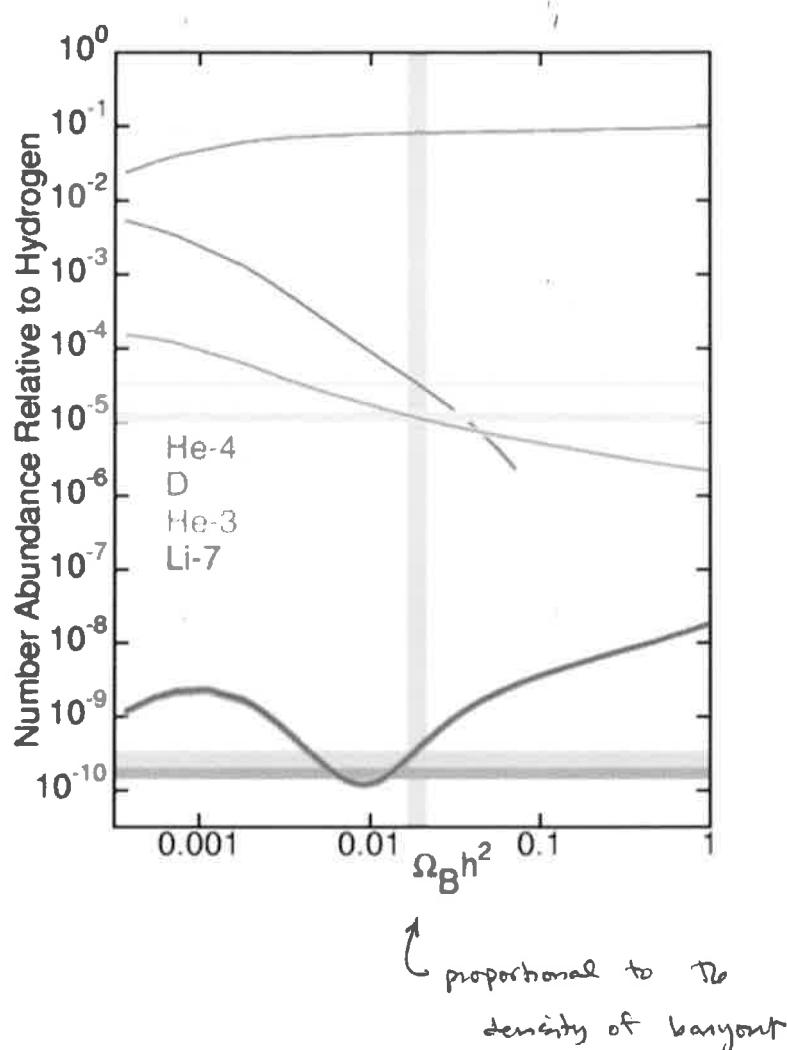
problems (or interesting features) of consensus model :

1. flatness (why is  $\Omega_m$  so close to 1? <sup>if it closer</sup> it was much closer in the past)
2. horizon problem: opposite sides of the sky : CMB has same properties but those 2 spots were never in physical contact w/ each other

also  $\rightarrow$  Sean Carroll's talk — why was early universe in a low entropy state?

proposed solution  $\rightarrow$  INFLATION — brief period of very fast, exponential expansion in the very early universe  $\rightarrow$  because of huge expansion, all of the observable universe was in causal contact; curved space gets very flat (due to stretching)

\* What about "fine-tuning"?  $\Rightarrow$   $H_0, \Omega_m, \Omega_b$  seem designed to give the universe complexity



proportional to  $\Omega_B$   
density of baryons

\* horizontal lines are observed  
values of each element