Proposal for Chandra Observations

Cover Page

| Principal Investigator Prof. David H Cohen | | |
|--|----------------------------|------------------------|
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| Proposal Title Quantitative X-ray Spectral Modeling of the Canonical O Supergiant Wind-Shock Source zeta Puppis | | | | | | | | | | |
|---|---------------------|------------------------|--------------------|--|--|--|--|--|--|--|
| Subject Category STARS AND WD | | | | | | | | | | |
| Proposal Type G0 | Linked Proposal | Distr. Medium WWW ONLY | Proprietary Rights | | | | | | | |
| Total Requested Time 210.00 | Number of Targets 1 | | Proposed Budget | | | | | | | |

| Joint Proposal? | |
|-----------------|-----------------------------|
| HST Orbits | HST Instruments: |
| XMM Time | RXTE Time: |
| NOAO Nights? | NOAO Telescope/Instruments: |
| NRAO Hours | NRAO Telescopes |

Abstract

The X-ray bright O supergiant zeta Puppis has a 68 ks archival HETGS spectrum that is rich with spectrally resolved diagnostic emission lines. The 210 ks observation we are proposing here will enable us, primarily by the quantitative modeling of the broad and asymmetric emission lines, to determine the properties of the shock-heated wind and to determine the relative roles of mass-loss rate reduction and large scale clumping and porosity. Not only will the proposed observation generate a spectrum that will be one of the most significant legacies of the Chandra gratings, but it will provide key information about the mass-loss rates of O stars, which will have a bearing on studies of the galactic ISM and the fates of evolved massive stars.

Proposal Number 10200630 Date: 2008-03-20 Admin. use only

${\bf Proposal~for~\it Chandra~\bf Observations}$

General Form

| PI Prof. David H Cohen |
|---|
| Proposal Title |
| Quantitative X-ray Spectral Modeling of the Canonical O Supergiant Wind-Shock |
| Source zeta Puppis |

| Co-Investigator(s) | | | | | | | | | |
|------------------------------|---|---------|--|--|--|--|--|--|--|
| First Name | E-Mail | | | | | | | | |
| Last Name | Institute | Country | | | | | | | |
| Maurice Leutenegger | maurice@milkyway.gsfc.nasa.gov NASA/GSFC | USA | | | | | | | |
| Janos Zsargo | jzsargo@bruno.phyast.pitt.edu UNIVERSITY OF PITTSBURGH | USA | | | | | | | |
| John Hillier | jdh@rosella.phyast.pitt.edu UNIVERSITY OF PITTSBURGH | USA | | | | | | | |
| Stan Owocki | owocki@bartol.udel.edu UNIVERSITY OF DELAWARE | USA | | | | | | | |
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| Are there additional Co-Is | listed in the science justification? | | | | | | | | |
| Is the first Co-I doing obse | rving, rather than the PI? N Telephone: | | | | | | | | |

Institute Endorsement

| Name of Administrator | Constance Hungerford |
|--------------------------|----------------------|
| Administrative Authority | Provost |
| Administrative Institute | Swarthmore College |
| Admin Signature: | Date: |
| PI Signature: | Date: |

Proposal for Chandra Observations

Target Summary

PI Prof. David H Cohen

Proposal Title
Quantitative X-ray Spectral Modeling of the Canonical O Supergiant Wind-Shock
Source zeta Puppis

| | bource zeta rappis | | | | | | | | |
|-----|---|-------------|------------|---------|-------------------|----------|------------|------------|----------|
| | Target Name | (J2000) | Offsets | | | Detector | (c/s) | | Grid |
| | Solar System Object | (====) | Y Detector | Optical | Observ. | Grating | Count Rate | Time- | |
| Tar | Grid Name | R.A. | Z Detector | Monitor | Time | HRC | 1st Order | Constr? | #Points |
| No | | Dec. | SIM Trans | V Meg | | Timing | Total Fld. | Constr. | MaxDist. |
| 1 | zeta Pup | 08 03 35.0 | SIM Hans | N - Mag | (ksec) 210.000 | ACIS-S | n 200000 | N EXU.SIC: | N N |
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Proposal for Chandra Observations

ACIS Parameters (Required, Pileup, Telemetry Parameters)

PI Prof. David H Cohen

Proposal Title

Quantitative X-ray Spectral Modeling of the Canonical O Supergiant Wind-Shock Source zeta Puppis

| | Exposure Mode | | CC | \mathbf{Ds} | On | | | Most Eff. | Suba | rray | Alte | rnating | | ergy Filter | Spect | ra |
|-----|----------------------|----|--------|---------------|--------|--------|----|--------------|--------------|---------------------|----------|-----------------------|-----|----------------------------------|--------------|----------------|
| Tar | Mode Telemetry. | | | | I2 | | | CCD | | StartRow | Exp | osures Nbr. Rows | | Lower <u>Thresh.</u> Range | Max | Mult. |
| No. | Telemetry. Format | S0 | S1 | S2 | S3 | S4 | S5 | Time | Type NONE | StartRow No.Rows | Y/N N | Nbr. Rows Exp.Time | Y/N | Range | Max Count | Mult. Lines |
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Proposal for Chandra Observations

ACIS Parameters (Custom:Telemetry Overflow Parameters)

PI Prof. David H Cohen

Proposal Title
Quantitative X-ray Spectral Modeling of the Canonical O Supergiant Wind-Shock
Source zeta Puppis

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| No der Chip Type Row Col Width Height Threshold Range Rate Additional Spatial Windows | |
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Proposal for Chandra Observations

Cycle 10

Target Constraints

| PI | Prof. | David H Cohen |
|-----|---------|--|
| Pro | posal T | tle |
| | Quantit | ative X-ray Spectral Modeling of the Canonical O Supergiant Wind-Shock |
| | Source | zeta Puppis |

| | | Window Constra | int | Roll Constraints | | | | | Phase Dependent Observations | | | | |
|-----------|------|----------------|-----------|------------------|------|-----------------|---------------------|------|------------------------------|------------------------|------------------------|--|--|
| Tar No | Flag | Start Time | Stop Time | Flag | 180? | Angle (degrees) | Tolerance (degrees) | Flag | Epoch(MJD) Period(days) | Min.Phase Min.Error | Max.Phase Max.Error | | |
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| | Monitoring Observations | | | | | Group Obser | vations | Un- | Coor | Add. | |
|-----------|-------------------------|---------------------|--|--|---|----------------------|---------|------|-----------------|------------------|--|
| Tar No | | Geometric Factor | | | | Interval D (days) | | Flag | Interval (days) | Con- straints | |
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Proposal for Chandra Observations

Cycle 10

TOO Details

| PI Prof. David H Cohen |
|---|
| Proposal Title |
| Quantitative X-ray Spectral Modeling of the Canonical O Supergiant Wind-Shock |
| Source zeta Puppis |
| |

| | | Alternates | | Response Window | | | | | Followup Observations | | | Obs.Params | |
|-----|-------|------------|------|-----------------|-------|------|---------|---------|-----------------------|------|--------|------------|--------------|
| Tar | Trig- | | Nbr. | | Q | α. | Prob- | Initial | | Obs. | | Tolerance | specified by |
| No | ger? | Group Name | Req. | (days) | Start | Stop | ability | Alloc. | Ordei | Time | (days) | (%) | Target No. |
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| TOO Trigger Criteria | |
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| TOO Followup Instructions | |
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If this TOO is a resubmission of a proposal approved in the previous Cycle, should this TOO be canceled if the previous Cycle TOO is triggered?

${\bf Proposal~for~\it Chandra~\bf Observations}$

Cycle 10

Target Remarks

| PI | Prof. David H Cohen |
|-----|---|
| Pro | oposal Title |
| | Quantitative X-ray Spectral Modeling of the Canonical O Supergiant Wind-Shock |
| | Source zeta Puppis |

| Tar No | Remarks Coordinated Observation: Observatories |
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