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esffrc.f

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c
c subroutine esffrc(dum,grav0,grav1)
c
c Ensemble local Source Function (ESF) method for radiation force.
c
c Version with "Piecewise Linear Source Function" (PLSF)
c
c Operation controlled by ifrc switch as follows:
c
c ifrc.le.0 => grad = gcak
c     .eq.1 => grad = gabs
c     .eq.2 => grad = gabs,      gscat=gssf
c     .eq.3 => grad = gabs+gscat, gscat=gssf
c     .eq.4 => grad = gabs+gscat, gscat=gesf w/ S(rp)=So(r)=const (~SSF har
d way)
c     .eq.5 => grad = gabs+gscat, gscat=gesf w/ S(rp)=So(rp)=Smooth VARIABLE
E source funct
c     .eq.6 => grad = gabs+gscat, gscat=gesf w/ S(rp)=So(rp)*betac(rp)/beta
(rp)+tauc=0.
c     .ge.7 => grad = gabs+gscat, gscat=gesf w/ S(rp)=So(rp)*betac(rp)/beta
(rp)+tauc.ne.0.
c
c Revision history:
c 11/24/97: made irssfmax max grid point for S=So in EISF model
c 11/24/97: set ifrc=7 for tauc.ne.0., retained ifrc=6 w/ tauc=0.
c 9/10/96: add tauc to eisf
c 2/7/95: tests of varying LBC vth (fvthlbc)
c 10/1/94: initial simulations using PLSF
c 9/25/94: initial simulations using PCSF
c 9/19/94: adapted from Smooth Source Function method routine ssffrc.
c 4/20/94: corrected to take account Feldmeier erratum treatment of vth(T(r))
c
c 7/15/91: limit frequency integration to local CMF frequency interval
c
c Method:
c Calculate ensemble absorption line RADiation FORce at nr radial grid pts,
c by integrating absorption in z along nray rays with impact parameters p,
c where 0 < p < R*, and z = sqrt(r**2-p**2).
c One-sided version of RADFRC for staggered mesh.
c Assumes linear variation of velocity and density between zone INTERFACES;
c Changes in optical depth data are then at ZONE-CENTERS, while eta,
c the optical depth itself, is at zone INTERFACES.
c The interface value of the radiative acceleration is then computed from
c grad = SUM(iray) SUM(ix) phi(x-vntf/vth)/eta**alpha.
c Note boundary condition on radiation force is at r=max(rstar,r0).
c Also,
c if (ifrc.gt.1), also computes SSF diffuse force,
c
c CMF frequency integrations limited to ixc+-nxcb2, where
c     ixc(ir)=ifix((v(r)/vth-x(1))/delx)+1
c     nxcb2=nxc/2=32
c Storage requirements minimized by computing inward ray optical
c depth eminr from detacmf, which spans only CMF x-range.
c
c include 'global.h'
c include 'sweep.h'
c include 'zone.h'
c integer stderr
c parameter (adumin=1.e-4,stderr=6)
c parameter (spinv=.5641895835)

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else
  yray=0.5
  wyray=1.
endif
c
c First, convert VH-1 to TDSW variables,
c taking TDSW interface = VH-1 zonal centers
c
  alpha = abbott
  nr   = nmax
  irad1 = nmin
  iradf = nmax
  rsscl = rstar
  oma   = 1.-alpha
  opa   = 1.+alpha
  alpham= -alpha
  obkapm= 1./xkapm
  obkmtoma=obkapm**oma
write(stderr,'*) 'obkapm=' ,obkapm
c
nrp1   = nr+1
irad0  = irad1-1
iradfp1= iradf+1
iprnt=irad0+50
c
do ir=irad1,iradf
  radz(ir)=xa0(ir)
enddo
do ir=irad0,iradf
  grad(ir) = fuz
  gscat(ir)= fuz
  gssf(ir) = fuz
  tauc(ir) = fuz
  vntf(ir) = u(ir)
  roi(ir)  = rho(ir)
  rntf(ir) = xa0(ir)+0.5*dx0(ir)
  asndz(ir)= sqrt(p(ir)/rho(ir))
enddo
rntf(irad0)=xa0(irad1)-0.5*dx0(irad1)
radz(irad0)=xa0(irad1)-dx0(irad1)
c
write(stderr,'*)
$ irado,vntf(irado),roi(irado),rntf(irado),asndz(irado)
asndre = sqrt(boltzman*tempw/avgmass)
vth   = vthba*asndre
bvth  = cak*vth**alpha
frnorm = spiinv*delx*oma
c
c Set range of freq grid.
c
  xvmin = xo
  xvrgn = xf-xo
  if (xf.lt.0.) then
    vmax = -1.e20
    vmin = 1.e20
    do i = 1 , nr
      vmax = max(vmax,vntf(i))
      vmin = min(vmin,vntf(i))
    enddo
    if (xf.lt.xo) then
      xvrgn = vmax/vth - 2.*xo
      xvmin = xo
    else

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      xvmax = vmax/vth-xf
      xvmin = max(xo,vmin/vth+xf)
      xvmin = delx*(int(xvmin/delx))
      xvrgn = xvmax-xvmin
    endif
  endif
1111  nx = 64*(int((xvrgn)/(delx*64.))+1)
c555  nx = (int((xvrgn)/(delx ))+1)
  nx = min(nxmax,nx)
  x = xvmin
  do ix=1,nx
    xvec(ix) = x
    x = x+delx
  enddo
c
c Begin RAY LOOP:
c
  do iy=1,nray
c
  c Store coefficients for x-integration:
  c
    do ir=irad0,iradf
      cost   = sqrt(1.-yray(iy)*(rsscl/rntf(ir))**2)
      irml  = max(irad0,ir-1)
      vthz  = vthba*asndz(ir)
      velm(ir) = cost*vntf(ir)/vthz
      zray(ir) = cost*rntf(ir)
      c4(ir)  = vth/vthz
      c5(ir)  = c4(ir)*exp(min(0.,1.-(asndz(ir)/asndre)**2))
      c4(ir)  = 1. ! No therm speed correction
      c5(ir)  = 1. ! No kappa temp. correction
      prad(ir)= wyray(iy)*c5(ir)/c4(ir)
    c Compute index for local CMF frequency,
    c and use to set x-integration range
      ixvec(ir)=ifix((velm(ir)/c4(ir)-xvec(1))/delx)+1
    enddo
    do ir=irad1,iradf
      irml  = ir-1
      rhoo  = roi(irml)
      dz    = zray(ir)-zray(irml)
      drho  = roi(ir)-rhoo
      c1(ir) = rhoo*dz
      c2(ir) = drho*dz
      c3(ir) = spiinv*c2(ir)/2.
    c
    if((ir.eq.iprnt).or.(ir.eq.iprnt+1))
    $ write(stderr,'*) ir,rho,dz,drho,c1(ir),c2(ir)
      ixavg = (ixcvec(ir)+ixcvec(irml))/2
      ixcdel = abs(ixcvec(ir)-ixcvec(irml))
      ixccfac = ifix(float(ixcdel)/nxcc+1.3)
      nxc   = nxcc
    c
      if (ixcdel.gt.(0.7*nxcc)) nxc=nxcc2
      nxc   = ((nxcdop+ixcdel)/nxcc+1)*nxcc
      nxc   = min(nx,nxc,nxcmax)
      ixcmnv(ir)=min(max(ixavg-nxc/2.0),nx-nxc)+1
      ixcmxv(ir)=ixcmnv(ir)+nxc-1
      if (nxc.gt.nxcc)
      c
      write(stderr,'*)ir,nxc,ixcmn,ixcmx,ixc,xvec(ixcvec(ir))-velm(ir)
    enddo
  c
  c Now initialize esumr at LBC
  c as Schuster-Schwarzschild reversing layer (for gabs):
  c

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do ix=1,nx
  xmu(ix) = xvec(ix)*fvthlbc*c4(irad0) - velm(irad0)
  ex(ix) = 0.
c   fc(ix) = cvmgp(0.,1.,xmu(ix))
c   fc(ix) = cvmgp(0.,1.,float(ix-ixcvec(irad0)))
  esumr(ix)= obkapt
enddo
ixcmin = ixcmnv(irad1)
ixcmax = ixcmxv(irad1)
ixcmnv(irad0) = ixcmin
ixcmxv(irad0) = ixcmax
nxc = ixcmax-ixcmin+1
write(stderr,'*) irad0,ixcmin,ixcmax,ixc,xmu(ixc)
call fcaphi(nxc,xmu(ixcmin),ex(ixcmin),fc(ixcmin))
call fcaphi(ixcmax,xmu,ex,fc,ixcmin)
fsum=0.
do ix=ixcmin,ixcmax
  esumr(ix) = xmo*spiinv*ex(ix) + obkapt
  tmp=ex(ix)/esumr(ix)**alpha
  fsum=fsum+tmp
enddo
grad(irad0) = grad(irad0)+prad(irad0)*fsum
do ix=1,nx
  elbc(ix)=esumr(ix)
enddo
c Begin RADIUS LOOP:
do ir=irad1,iradf
c Initialize x arrays over the full x-range by presetting
c the exponential and error function arrays
c to their limiting forms:
do ix=1,nx
  exo(ix) = ex(ix)
  fco(ix) = fc(ix)
  ex(ix) = 0.
c   fc(ix) = cvmgp(0.,1.,ix-ixcvec(ir))
c   fc(ix) = cvmgp(0.,1.,float(ix-ixcvec(ir)))
enddo
c Now over the limited CMF x-range, perform the following:
c First, compute exponential & error functions ex(ix) & fc(ix),
c and use these to obtain esumr(ix) increments deta.
c Then compute force contribution by summing
c deta/esumr(ix)**alpha
c Note that because of possible x-range differences between
c zones ir and ir-1, it is necessary to recompute xmuo(ix).
c However, exo(ix) and fco(ix) are known over the full x-range
c by virtue of the preset to their asymptotic values.
c
  irm1 = ir-1
  ixcmin = ixcmnv(ir)
  ixcmax = ixcmxv(ir)
  nxc = ixcmax-ixcmin+1
  do ix=ixcmin,ixcmax
    xmu(ix) = xvec(ix)*c4(ir) - velm(ir)
    xmuo(ix) = xvec(ix)*c4(irm1) - velm(irm1)
  enddo

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c   call fcaphi(nxc,xmu(ixcmin),ex(ixcmin),fc(ixcmin))
c   call fcaphi(ixcmax,xmu,ex,fc,ixcmin)
fsum = 0.
do ix=ixcmin,ixcmax
  dxmu = xmuo(ix)-xmu(ix)
  dxmusq= dxmu*dxmu
  if (dxmusq.lt.fuz) then
    deta = c1(ir)*(ex(ix)+exo(ix))
  else
    deta = ((c1(ir)*dxmu+c2(ir)*xmuo(ix))*(
              (fc(ix)-fco(ix))-
              c3(ir)*(ex(ix)-exo(ix)))/dxmusq
  endif
  if((ir.eq.iprnt).or.(ir.eq.iprnt+1)).and.(ix.eq.9))
  write(stderr,'*) ix,ir,
  (c1(ir)*dxmu+c2(ir)*xmuo(ix))*(fc(ix)-fco(ix)),
  c1(ir)*dxmu,c2(ir)*xmuo(ix),fc(ix)-fco(ix),
  -c3(ir)*(ex(ix)-exo(ix)),dxmusq,deta
  deta = c5(ir)*max(deta,fuz) !kap temp corr
  esumr(ix) = esumr(ix) + deta
  ixc = ix-ixcmin+1
  detacmf(ixc,ir) = deta
  exc(ixc,ir) = ex(ix)
  tmp = ex(ix)/esumr(ix)**alpha
  if ((ix.eq.ixcmin).or.(ix.eq.ixcmax)) tmp=0.5*tmp
  fsum = fsum + tmp
enddo
c2(ir) = fsum !Store to vectorize normalization
c -- NEW~~~~~ JS -- Simplify later!
rsbr = rsscl/radz(ir) !make So ZONE-centered
rsbrsq = rsbr*rsbr
xmustar = sqrt(max((1.-rsbrsq),0.))
cost = sqrt(1.-yray(iy)*rsbrsq)
f_ld(ir) = (cost**2-xmustar**2)/(1.-xmustar**2)
if (f_ld(ir).lt.0.0) stop 'ld corr, dir<0'
b_mu(ir) = fsum*sqrt(f_ld(ir))
write(*,*) cost,xmustar,sqrt(f_ld(ir))
enddo ! End Main Radius loop for direct term
do ir=irad1,iradf
  if (epsabs.le.-1.0) then
    c ld correction, JS0412
    grad(ir) = grad(ir) + 0.5*prad(ir)*c2(ir) + (3./4.)*prad(ir)*b_mu(ir)
    write(*,44) ir,radz(ir)/rsscl,c2(ir),grad(ir),b_mu(ir),sqrt(f_ld(ir))
    c OBS! This is actually a *constant* as it is now -- increase force by ~3% for y=0.5 ray
    else
      grad(ir) = grad(ir) + prad(ir)*c2(ir)
    endif
  enddo
  FORMAT(i10,5e15.5)
c Now conditionally compute SSF approx diffuse force.
c   if (ifrc.gt.1) then
c     First implement eta_minus RBC.
c     For ISSFRBC <=0, use SPO RBC:
c     Assume eta_minus(RBC) is ~ rho*r
c     (i.e., velocity constant, density ~1/r^2) ---

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c      this zeroes out ftot.
c
c      fsum=0.
do ix=1,nx
  xmu(ix) = xvec(ix)*c4(iradf) - velm(iradf)
enddo
if (issfrbc.le.0) then
  do ix=1,nx
    tmp = xmu(ix)
    tmp = spiinv*exp(-tmp*tmp)
    eminr(ix) = roi(iradf)*rntf(iradf)*tmp+obkapm
  enddo
c For ISSFRBC >=1, use JIC RBC:
c Assume eta_minus is eta_plus reflected about CMF line center.
c (i.e., Odd symmetry of v wrt rmax, and even symmetry of rho) ---
c this zeroes out gscat.
c
else
  do ix=1,nx
    xref = -xmu(ix)
    uref = (xref-xmu(1))/delx+10001.
    ixr = int(uref) - 10000.
    uref = uref - 10000. - float(ixr)
    if ((ixr.le.0).or.(ixr.ge.nx)) then
      eminr(ix) = obkapm
    else
      eminr(ix) = esumr(ix)+uref*(esumr(ixr+1)-esumr(ixr))
    endif
  $      write(stderr,375) ix,ixr,xref,uref,esumr(ix),esumr(ixr),
c      $          eminr(ix)
375  format(' RBC:',2i4,1p6e10.3)
    enddo
  endif
  do ix=1,nx
    erbc(ix)=eminr(ix) ! save rbc for later
  enddo
c Now do BACKward spatial integral for the
c inward contribution to diffuse force.
c
  xkapc = elkap
c  xkapc = xkapc+1./c1(irad1)
c **BUG**  tauc(iradf) = xkapc*rho(iradf)*radz(iradf)
tauc(iradfp1) = 0.
do ir=irad1,irad1,-1
  tauc(ir) = tauc(ir+1)+xkapc*c1(ir)
  write(stderr,'*) ir,c1(ir),xkapc,tauc(ir)
  ixcmn = ixcmnv(ir) !ir-1 better???
  ixcmx = ixcmxv(ir) !
  fsum = 0.
  do ix=ixcmn,ixcmx
    ixc = ix-ixcmn+1
    tmp = exc(ixc,ir)
    tmp = tmp/minr(ix)**alpha
    if ((ix.eq.ixcmn).or.(ix.eq.ixcmx)) tmp=0.5*tmp
    fsum = fsum + tmp
    eminr(ix) = eminr(ix)+detacmf(ixc,ir)
  enddo
  c2(ir) = fsum
enddo ! end SSF radius loop

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do ir=irad1,irad1
  gscat(ir) = gscat(ir) + prad(ir)*c2(ir)
enddo
endif ! SSF
enddo ! end of ray loop
c Next, conditionally compute smooth source function, sobicrr.
c
if (ifrc.gt.1) then
  do ir=irad0,irad1
    rsbr   = rsscl/rntf(ir) !make So INTF-centered
    rsbr   = rsscl/radz(ir) !make So ZONE-centered
    rsbrsq = rsbr*rsbr
    xmstar = sqrt(max((1.-rsbrsq),0.))
    thinfac = 0.5/(1.+xmstar)
    sobicrr(ir) = zso(max(ir-2,1),1,1)
    if (sobicrr(ir).le.0.) sobicrr(ir)=thinfac !make sure So defined.
c Compute Source function correction for given eps assuming Bplanck/Ic = 1.
    if (epsabs.lt.0.) then
      tmp = thinfac-epsabs/rsbrsq
      sobicrr(ir) = tmp/(1.-epsabs)                                !set So= opt. THIN form
      if (epsabs.le.-1.) then
        call s_limb_dark(sobi_ld,xmstar)
        sobicrr(ir) = sobi_ld/rsbrsq
c HAVE ADDED TEST_CASES FOR LIMB-DARK HERE
c -- Don't forget the r^2 correction-factor!!
      endif
      write(*,*) 1./rsbr, sobicrr(ir),thinfac
      if (epsabs.le.-10.) sobicrr(ir) = -epsabs-1. !set So= const.
      else if (epsabs.gt.0) then                                    !set So= opt. THICK form
        (OR-II modified)
        dvbdr= (vzone(ir)-vzone(ir-1))/(radn(ir)-radn(ir-1))
        dvbdr= (vntf(ir)-vntf(ir-1))/(rntf(ir)-rntf(ir-1))
c Abs value is poor man's way of dealing with nonmonotonicity; improve later.
        dvbdr= abs(dvbdr)
        vbr = abs(vntf(ir)/rntf(ir))
        sig = dvbdr/flr(vbr)-1.
        vbr = vinfsc1*(1.-rsbr)/rntf(ir) ! use smooth beta=1 vel. law
        sig = (2.*rsbr-1.)/flr(1.-rsbr)
        e2 = 1.+xmstar
        e3 = 1.+e2*xmstar
        e4 = 1.+e3*xmstar
        e5 = (1.+e4*xmstar)/5.
        e4 = e4/4.
        e3 = e3/3.
        e2 = e2/2.
        qc = 1.+sig*e3
        q = 1.+sig/3.
        if (sig.lt.-1.) then                                         !Do proper correction for dv/dr<0.
          xo = sqrt(-1./sig)
          eo3= (1.+xo*(1.+xo))/3.
          cf = (1.-xo)*(1.+sig*eo3)
          q = q - cf
          xo = max(xo,xmstar)
          eo3= (1.+xo*(1.+xo))/3.
          cf = (1.-xo)*(1.+sig*eo3)
          cf = cf*2./(1.-xmstar)
          qc = qc-cf
        endif
        et = (epsabs*epscl*roi(ir)**2)/vbr !eps'*tauo
        tmp = thinfac*qc + et/rsbrsq
        sobs= tmp/(q+et)
      endif
    endif
  enddo
enddo

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sobicrr(ir) = sobs
c Now correct to get right asymptotic growth rates.
c     sig= abs(dvbldr)/flr(vbr)-1. !revert to poor man's method here.
    if (sig.gt.-1.) then
        tmp = e3+sig*e5-4.*e2*(1./3.+sig/5.)*sobs
        tmp = 0.5*(1.-tmp/(e2+sig*e4))
        sobicrr(ir) = tmp
    endif !** endif sig.gt.-1
    endif !** endif epsabs.gt.0
    enddo
endif !** endif ifrc.gt.1 => Compute sobicrr
c If not ESF, normalize now.
c
if (ifrc.lt.4) then
    do ir=irad1,iradf
        tmp = frnorm*bvth/(rntf(ir)*rntf(ir))
        gscat(ir) = tmp*(gscat(ir)-grad(ir))*sobicrr(ir)
        gssf(ir) = gscat(ir)
        grad(ir) = tmp*grad(ir)
        if (ifrc.eq.3) grad(ir)=grad(ir)+gscat(ir)
    write(stderr,*) ir,gscat(ir),grad(ir)
    enddo
else
c local, Ensemble Source Function (ESF) force option:
c (** NOTE: Currently assumes only a single ray ***)
c
c First, build ensemble source function into c3(ir).
c
    do ir=irad1,iradf
        tmp=2./(1.+(gscat(ir)+gscat(ir-1))/(grad(ir)+grad(ir-1)))
        if (ifrc.eq.5) tmp=1. ! Smooth VARIABLE
Source Func.
        if ((ir.lt.irad0+10).and.notinit) tmp=zef(ir-2,1,1) ! Use init S n
ear LBC.
        if ((ir.lt.irad0+irssfmax)) tmp=1. ! Test to keep LBC
smooth.
        if ((ir.lt.irad0+irssfmax)) tmp=0. ! Test to destab.
base
        sobic=(sobicrr(ir)+sobicrr(ir-1))/(2.*radz(ir)*radz(ir)) ! zone-cen
ter So
        c3(ir)=(tmp+2.*ih(ifrc-6)*tauc(ir))*sobic ! S(r), including
tauc for ifrc>6
        if (ifrc.eq.4) c3(ir)=1. ! SSF the hard way
, for testing.
        write(stderr,990) ir,sobicrr(ir),c3(ir),tmp,gscat(ir),grad(ir)
    enddo
notinit=.true.
c
c For both forward and backward rays, compute nested rp sum of intensity differ
ential,
c and then frequency integrate for forward-backward stream intensity.
c
    do ir=irad0,iradf
        do ix=1,nx ! Reinitialize for calculation of eta(r)-
eta(rp)
            esumr(ix)= elbc(ix)
            eminr(ix)= erbc(ix)
            esumr(ix)= obkpm
c

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c
eminr(ix)= obkpm
fc(ix) = 0.
xmu(ix) = 0.
xmuo(ix) = xmu(ix)
enddo
if (ir.gt.irad0) then
    do irp=ir,irad1,-1 !forward ray
        ixmin = max(ixcmnv(ir),ixcmnv(irp))
        ixmax = min(ixcmxv(ir),ixcmxv(irp))
        do ix = ixmin,ixmax
            ixcrp = ix-ixcmnv(irp)+1
            esumr(ix) = esumr(ix)+detacmf(ixcrp,irp)
            etma = esumr(ix)**alphan
            tmp=c3(irp)*(xmu(ix)-etma)
            fc(ix)=fc(ix)+tmp
            xmu(ix)=etma
        if ((ir.eq.ipnrt).and.(ixcrp.ge.29).and.(ixcrp.le.35)
            .and.(irp.ge.40))
            if ((ir.eq.iprnt).and.(ix.eq.ixcvec(iprnt))) then
                write(stderr,980)
$ ix,ixcrp,irp,c3(irp),detacmf(ixcrp,irp),esumr(ix),tmp,fc(ix)
            endif
980    format(1x,3i5,1p6e11.4)
        enddo
    enddo
endif
if (ir.lt.irad0) then
    do irp=ir+1,iradf !backward ray
        ixmin = max(ixcmnv(ir),ixcmnv(irp))
        ixmax = min(ixcmxv(ir),ixcmxv(irp))
        do ix = ixmin,ixmax
            ixcrp = ix-ixcmnv(irp)+1
            eminr(ix) = eminr(ix)+detacmf(ixcrp,irp)
            etma = eminr(ix)**alphan
            tmp=c3(irp)*(xmuo(ix)-etma)
            fc(ix)=fc(ix)-tmp
            xmuo(ix)=etma
        if ((ir.eq.iprnt).and.(ixcrp.ge.29).and.(ixcrp.le.35)
            .and.(irp.le.61))
            if ((ir.eq.iprnt).and.(ix.eq.ixcvec(iprnt))) then
                write(stderr,980)
$ ix,ixcrp,irp,c3(irp),detacmf(ixcrp,irp),eminr(ix),tmp,fc(ix)
            endif
        enddo
    enddo
endif
fsum = 0.
irp1=min(ir+1,iradf)
do ix = ixcmnv(ir),ixcmxv(ir)
    ixcr = ix - ixcmnv(ir) + 1
    ixcr1= ix - ixcmnv(irp1) + 1
    ixcr1= min(max(1,ixcr1),ixcmxv(irp1))
    tmp = flr(0.25*(detacmf(ixcr,ir)+detacmf(ixcr1,irp1)))
    tmp = detacmf(ixcr,ir)
    if((ix.ge.ixcmnv(irp1)).and.(ix.le.ixcmxv(irp1))) then
        tmp=tmp+detacmf(ixcr1,irp1)
        tmp=flr(0.25*tmp)
        tmp2= 1.
ng)
c
tmp2= ((obkpm+tmp)**oma-obkmtoma)/(oma*tmp) !PLSF
fc(ix) = fc(ix)+tmp2*(c3(ir)-c3(ir+1)) !local S' term corre
ction

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tmp = exc(ixcr,ir)*fc(ix)
if ((ix.eq.ixcmnv(ir)).or.(ix.eq.ixcmxv(ir))) tmp=tmp/2.
fsum = fsum + tmp
if ((ir.eq.iprnt).and.(abs(ix-ixcvec(ir)).le.5)) then
tmp3=1./eminr(ix)**alpha-1./esumr(ix)**alpha
if (ifrc.ne.4) tmp3=tmp3*sobicrr(ir)/(rntf(ir)*rntf(ir))
write(stderr,990)
$ ix,eminr(ix),esumr(ix),fc(ix),tmp2,tmp,fsum,exc(ixcr,ir)
endif
enddo
c2(ir) = fsum           ! ~ Ibar^(+r)-Ibar^-(r)
enddo
tmp1=grad (irad1)
tmp2=gscat(iradf)
if (ifrc.ne.4) then
tmp1=tmp1*sobicrr(irad0)/(rntf(irad0)*rntf(irad0))
tmp2=tmp2*sobicrr(iradf)/(rntf(iradf)*rntf(iradf))
endif
c2(irad0) = c2(irad0)+tmp1
c2(iradf) = c2(iradf)-tmp2

c Normalize to forces.

do ir=irad0,iradf
rsq = rntf(ir)*rntf(ir)
sobic= sobicrr(ir)/rsq
if (ifrc.eq.4) c2(ir)= c2(ir)*sobic
tmp = frnorm*bvth
gssf(ir) = tmp*(gscat(ir)-grad(ir))*sobic
if(ifrc.gt.6)
grad(ir)=(tauc(ir)+1.)*grad(ir)-tauc(ir)*gscat(ir) ! tauc correcti
$ on.
grad(ir) = tmp*grad(ir)/rsq
gscat(ir) = tmp*c2(ir)
if(ir.le.irad0+10) gscat(ir)=gssf(ir)      *** TEST using SSF near L
900 write(stderr,990)ir,sobic,c3(ir)/sobic,grad(ir),gssf(ir),gscat(ir)
format(1x,i5,1p7e12.4)
enddo
gscat(irad0)=gssf(irad0)
gscat(iradf)=gssf(iradf) ! Revert to SSF for RBC.
do ir=irad0,iradf
grad(ir) = grad(ir)+gscat(ir)
enddo
endif !** endif ifrc.lt.4 , else

c Store results for VH-1

1000 continue
do n = nmin, nmax
c Option to turn off rad force if v-del below some negative cut-off
JS 1204
if (neg_frad.eq.1) then
delv = u(n)-u(n-1)
red_fac = exp(min(0.0,delv)/delv_cut)
grav0(n) = grav0(n) + grad(n)*red_fac
grav1(n) = grav1(n) + grad(n)*red_fac
if (red_fac.lt.1.0) write(*,*) n,delv,grav0(n),grad(n),exp(min(0.0,
delv)/delv_cut),delv_cut
else
c Here comes standard option

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t = 1./(1.+p*abs(x(ix)))
fc(ix) = t*(a1+t*(a2+t*a3))*ex(ix)
enddo
c   if (xvoigt.gt.0.) then
c     do ix = 1 , nx
c       ax      = flr(abs(x(ix)))
c       xvmx   = xvoigt-ax
c       ev      = avoigt/(ax**2)
c       ex(ix) = cvmgp(ex(ix),ev,xvmx)
c       ex(ix) = ex(ix)*bvoigt
c       fv      = cvoigt/ax
c       fc(ix) = fc(ix) + dvoigt
c       fc(ix) = cvmgp(fc(ix),fv,xvmx)
c       fc(ix) = fc(ix)*bvoigt
c     enddo
c   endif
c   do ix = 1 , nx
c     fc(ix) = cvmgp(fc(ix),1.-fc(ix),x(ix))
c   enddo
c write(stderr,10) x,ex,t,fc
c10 format(' fcaphi:'1p5e10.3)
c return
c
c subroutine s_limb_dark(s_ld,mustar)
c
c Include Eddington limb-darkening
c in the optically thin scattering source function
c
c Use approximation below mustar=1e-3
real s_ld,mustar,sinmustar
if (mustar.gt.1.0.or.mustar.lt.0.0) then
  write(*,*),'mustar:',mustar
  stop'0>mustar or 1<mustar'
else if (mustar.gt.1.e-3) then
  sinmustar = sqrt(1.-mustar**2.)
  s_ld = (1./16.)*(7.-4.*mustar+3.*mustar**2.*alog(mustar/(1.+sinmustar))
/sinmustar)
else
  s_ld = (1./16.)*(7.-4.*mustar-3.*mustar**2.)
endif
return
end

subroutine angle_weights(yray,wyray,nray)

c Set up angle integration steps (y=[0,1]) and weights
c Use simple Trapez for testing for now
integer nray, i
dimension yray(nray),wyray(nray)
real dyray

dyray = 1. / (nray - 1)
do i=1,nray
  yray(i) = (i-1)*dyray
  if (i.eq.1.or.i.eq.nray) then
    wyray(i)=0.5*dyray
  else
    wyray(i)=1.0*dyray
  endif
  write(*,*),i,yray(i),wyray(i)
c

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enddo
stop
return
end

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